1 Magma-shale interaction in large igneous provinces: implications

2 for climate warming and sulfide genesis

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- 24 Running title: Magma-shale interaction in LIPs

- 26 Manuscript for submission to *Journal of Petrology* as a *Letter*.
- 27 For potential inclusion in the Themed Article Collections Large Igneous Provinces,
- 28 Experimental Petrology, and Metamorphic Reactions and Metamorphic Fluids.

29 ABSTRACT

30 Large Igneous Provinces (LIPs) whose magma plumbing systems intersect sedimentary basins are linked to upheavals of Earth's carbon and sulfur cycles and thus climate and life 31 32 history. However, the underlying mechanistic links between these phenomena are elusive. We address this knowledge gap through short time-scale petrological experiments (1200 °C 33 and 150 MPa) that explore interaction between basaltic melt and carbonaceous shale 34 (mudstone) using starting materials from the Canadian High Arctic LIP and the Sverdrup 35 36 Basin in which it intrudes. Here we show that entrainment of shale xenoliths in basaltic melt 37 causes shale to shatter due to incipient thermal stress and devolatilization, which accelerates assimilation by increasing reactive surface area. Shale assimilation therefore 38 facilitates transfer of sediment-derived volatile elements to the shallow parts of LIP 39 plumbing systems, whereupon carbon dominates the vapor phase whilst sulfur is 40 partitioned into sulfide melt droplets. This study reveals that although carbon and sulfur are 41 efficiently mobilized as a consequence of shale assimilation, sulfides can sequester sulfur -42 43 an important climate cooling agent - thus enhancing net emissions of climate warming 44 greenhouse gases by shale-intersecting LIPs.

45 **Abstract word count:** 180 words.

Keywords: C cycle perturbations; High Arctic LIP (HALIP); large igneous provinces; magmashale interaction; sulfide genesis

48 Main text: 3555 words + figure captions: 462 words = total word count: 4017

49 INTRODUCTION

50 Large Igneous Provinces (LIPs) are comprised of massive emplacements of largely basaltic magma into Earth's lithosphere and onto the surface (Bryan and Ernst, 2008; Black et al., 2021). Many LIP 51 52 events coincide with environmental change and mass extinctions (Courtillot and Renne, 2003; Bond 53 and Grasby, 2017; Ernst and Youbi, 2017), but the mechanisms underpinning this correlation are not 54 fully understood (Bond and Grasby, 2017; Black et al., 2021). The magnitude of carbon and sulfur 55 released to the atmosphere is generally accepted as a key control on the severity of environmental 56 impact of a LIP event (Jones et al., 2015). However, there is active debate concerning whether the 57 release of primary, mantle-derived volatiles by LIPs is sufficient to cause the global carbon cycle 58 disruptions, climate change, and biotic crises that mark mass extinctions; or whether additional 59 release of thermogenic, sediment-derived volatiles is required (Svensen et al., 2004, 2007, 2009; 60 Ganino and Arndt, 2009; Yallup et al., 2013; Callegaro et al., 2021; Capriolo et al., 2021, 2022; 61 Heimdal *et al.*, 2021).

62 It is therefore necessary to clarify the mechanisms, dynamics, and efficiency of volatile 63 removal from sedimentary rocks that are intersected by LIP plumbing systems. However, study of 64 magma-sediment interaction involving volatile-rich sedimentary rocks is fraught with obstacles such as poor outcrop preservation and accessibility. Sill-host rock contacts are often difficult to access and 65 66 deeply weathered, sometimes necessitating collection of drill cores such as for the Siberian LIP 67 (Callegaro et al., 2021) and the Brazilian portion of Central Atlantic Magmatic Province (Heimdal et al., 2019). Geochemical modelling of magma-sediment interaction is also limited, because whilst 68 69 several platforms exist for modelling assimilation of silicate rocks, thermodynamic data for non-70 silicate materials, such as carbonates, coal, organic matter, and evaporitic sulfates and halides are 71 lacking (Heinonen et al., 2021). Efforts to study magma-sediment interaction through direct 72 experimentation (using high pressure-temperature devices) has potential for generating unique 73 mechanistic insights but has thus far largely been restricted to studies involving magmatic 74 assimilation of carbonates (Freda et al., 1997, 2008, 2010; Iacono Marziano et al., 2007, 2008; Mollo et al., 2010; Deegan et al., 2010, 2016; Jolis et al., 2013; Carter and Dasgupta, 2015, 2016, 2018), 75 76 although there is growing interest in non-carbonate lithologies too (e.g. lacono-Marziano et al., 77 2017). Many of these previous studies utilized relatively long-duration experiments, with run times 78 ranging from several hours to days, which yielded equilibrium melt-gas assemblages that can be 79 difficult to interpret in terms of process.

80 Short-duration (seconds to minutes) petrological experiments that better replicate the onset 81 of magma-sediment interaction emerged in the early 2010s (Deegan *et al.*, 2010) and paved the way 82 for close examination of the rapid transformations that occur when basaltic melt first comes into 83 contact with sedimentary rock. In this paper, we apply a similar philosophy to simulate entrainment 84 of carbonaceous mudstone (hereafter referred to as "shale") xenoliths in basaltic melt and explore 85 incipient magma-shale interaction at LIPs. The experiments preserve textural and chemical 86 information that is largely obliterated in longer duration experiments as well as in many natural 87 contexts, which rarely freeze-in prograde reaction paths (Fig. 1a). Magma-shale interaction 88 moreover involves a sulfur component, which extends its relevance to sulfide ore formation. It is 89 generally accepted that sulfide formation in LIP-hosted Ni-Cu-PGE (Platinum Group Element) 90 deposits often involved incorporation of sedimentary crustal sulfur during magma emplacement 91 (Lesher, 2019). Since the solubility of sulfide in silicate melts is low, most of the sulfur needed to 92 constitute Ni-Cu-PGE ore bodies calls for addition of crustal sulfur as sulfide "xenomelt(s)", i.e., a 93 foreign melt derived from crustal rocks or xenoliths (Lesher, 2017, 2019). The crustal origin of sulfur 94 is supported by isotopic and geochemical data but uncertainty remains surrounding how sulfide 95 xenomelts form and are transported in ore-forming systems (Hayes et al., 2015; Lesher, 2017, 2019). 96 These are questions that short-duration high pressure-temperature interaction experiments can also 97 potentially clarify.

98 FIELD AREA, STARTING MATERIALS, AND METHODS

99 Here we employ the Cretaceous High Arctic LIP (HALIP) as a case study because its mafic sill 100 province, representing the frozen remains of the conduit system formerly supplying flood basalts to 101 the surface, is spectacularly preserved and exposed in the Canadian Arctic (Jowitt et al., 2014; 102 Evenchick et al., 2015; Deegan et al., 2018; Bédard et al., 2021a, 2021b). Much of the Canadian 103 portion of the HALIP was emplaced at ca. 120 Ma into the Sverdrup Basin, a sedimentary depocenter 104 filled with up to 13 km of siliciclastic, evaporitic, and carbonaceous strata of Carboniferous to 105 Paleogene age (Embry and Beauchamp, 2019). HALIP sills that invaded Sverdrup Basin sedimentary 106 rocks collectively influenced basin-scale thermal evolution and development of regional oil and gas 107 resources (Jones et al., 2007; Goodarzi et al., 2019) and possibly also sulfide ore formation (Jowitt et 108 al., 2014; Saumur et al., 2016). The Middle Triassic Murray Harbour Formation forms part of the 109 Sverdrup Basin fill and is dominated by black shale and siltstone, with subordinate calcareous and phosphatic interbeds. This unit is rich in Type II kerogen and is considered to be the primary source 110 111 rock for hydrocarbon discoveries in the Sverdrup Basin (Brooks et al., 1992; Kondla et al., 2015).

We utilized natural HALIP magmatic rock and Murray Harbour Formation shale as the experimental starting materials. The experimental conditions and compositions of the starting materials are provided in the **Supporting Data File**. Details of the experimental method and

analytical techniques applied to the experimental products are provided in full in the Supporting 115 116 **Information**. To summarise, the magmatic starting material is a pristine, non-cumulate mafic rock 117 sample from a ca. 30 m thick sill with 6.6 wt.% MgO and 0.15 wt.% S. The shale is a finely clastic 118 mudstone collected from a site over 60 m perpendicular distance away from any observable sill 119 contacts and contains 4.7 wt.% total organic carbon and 0.5 wt.% total S. The shale is principally 120 composed of quartz (56 wt.%), calcite (24 wt.%), biotite (9 wt.%), dolomite (4 wt.%), and pyrite (1 wt.%) and has a loss on ignition (LOI) value of 21.6 wt.%, indicating that roughly one fifth of its mass 121 comprises volatile compounds including sulfur from pyrite and organic matter at 0.5 wt.% plus 122 123 inorganic and organic carbon.

124 The magmatic rock was converted to glass, both an anhydrous and a mildly hydrated aliquot, 125 pulverized in an agate mill, and loaded into platinum capsules (ca. 25 to 35 mg of powdered glass 126 per experiment) along with a solid fragment of shale (ca. 4 to 5 mg per experiment). Each capsule 127 was positioned in a 19-25 mm NaCl-crushable MgO-borosilicate glass assembly employing a graphite 128 furnace and was then pressurized to 150 MPa in a low-P calibrated non-end loaded piston cylinder 129 (PC) device (see Masotta et al., 2012). The assemblies were then heated to 1200 °C at a rate of 100 130 °C per minute and held at the set point temperature for durations of 0 s, 300 s, and 600 s while 131 being continuously monitored for any P-T fluctuations. The experimental P-T conditions are 132 considered appropriate for injection of HALIP tholeiites at or slightly above their liquidus (1100 to 133 1200 °C) at ca. 5 km depth in the Sverdrup Basin (see Bédard et al., 2021a). The experiments were 134 terminated by shutting down the power source whereafter they were isobarically quenched at a rate 135 of ca. 100 °C per second in the first 5 seconds (to the glass transition) and ca. 33 °C per second 136 thereafter. Fast, isobaric quenching is important to preserve textures and glass (≈melt) that formed 137 in-situ within the time window of xenolith dissolution and melting (Fig. 1a). After quenching, the 138 capsules were retrieved, cast in epoxy, polished, and inspected using SEM (Secondary Electron 139 Microscopy). Quantitative analysis was subsequently performed utilising EPMA (Electron Probe 140 Microanalysis) and Raman micro-spectroscopy for glass and volatile compositions, respectively (see 141 **Supporting Information** for analytical details).

142 EXPERIMENTAL RESULTS

An extended description of the experimental products is provided in the Supporting Information
 and the compositions of the glasses and sulfides in the experimental products are provided in the
 Supplementary Data File. An overview of the textures of the experimental products is presented in
 Fig. 1 as a series of SEM mosaic images and a summary of the compositions of the experimental

glasses, sulfides, and vapour bubbles are provided in Figs. 2 and 3. In brief, the experiments
document features of incipient magma-shale interaction, including the following processes:

- 149 (i) Dissolution of shale into the host melt, manifest as disequilibrium textures and crustally
 150 contaminated, "modified" melts (Figs. 1 and 2).
- 151 (ii) Formation of modified melts that are initially restricted to patches within the shale
 152 fragment but eventually form a boundary layer surrounding shale. Calcium is enriched
 153 (relative to the starting HALIP basalt) in modified glass in all experiments whilst silica,
 154 potassium, and sulfur are often enriched too (Fig. 2).
- 155 (iii) Shattering and degassing of shale and formation of abundant vapor bubbles containing
 156 various carbon volatile species such as CO and CO₂ (Fig. 3).
- 157 (iv) Formation of sulfide mineralizations (≈ sulfide xenomelts) at magma-shale interfaces
 158 (Figs. 2 and 3).
- The experiments therefore reveal that contaminated melts and carbon volatiles are generated rapidly as a consequence of shale assimilation but that sulfides effectively sequester sulfur, a climate cooling agent, thus enhancing net emissions of greenhouse gases by LIPs that intersect shale-bearing sedimentary basins (as discussed below).

163 DISCUSSION

164 Magma-sediment interaction in microcosm

165 Our experiments have replicated magma-shale interaction in microcosm, i.e., the experiments

166 encapsulate in miniature the characteristics of a much larger system, as discussed here. Earlier,

similarly designed experiments showed that magma-carbonate interaction results in a low viscosity,

168 Ca-rich compositional boundary layer at the reaction site and a voluminous C-O-H (carbon-oxygen-

hydrogen) vapor phase (e.g. Deegan *et al.*, 2010). We find evidence of this process in our

170 experiments, but shale assimilation is complicated by the fact that shale contains silicate minerals

and organic matter in addition to carbonate. Previous experimental work involving siliceous shale

- and/or clay showed that at low pressure, clay begins to melt at ca. 1200 °C, but disequilibrium
- 173 melting of micas, quartz, and feldspar begins at ca. 700 to 800 °C, producing Si-rich partial melts
- 174 (Wyllie and Tuttle, 1961; Nichols *et al.*, 1996; Erdmann *et al.*, 2007). Heating experiments have also
- 175 recently shown that C-O-H-S fluids are generated from shale at temperatures ≤700 °C and that
- 176 silicate melt is produced from ca. 800 °C (Virtanen et al., 2021). None of these previous studies

simulated interaction between shale and basaltic melt, but there are nevertheless some broad
similarities to our work. For instance, two of our experiments contain a separated high-silica glass
phase (Fig. 1b,c), which we suggest is the product of quartz breakdown at <1200 °C during
experiment heat-up. These segregated high-silica glass domains are not present in the mildly
hydrated experiments, which suggests that melt mixing was more efficient at low temperatures
(during heat-up) when there was water added to the basaltic melt.

183 Importantly, we observed a portion of Ca-enriched, compositionally modified 184 (contaminated) glass near the reaction site in all of our magma-shale interaction experiments (Fig. 185 **2a,b**). The at times extremely high CaO content of this glass (45 wt.% max) indicates that the 186 carbonate components of the shale dissolved rapidly into the host melt. In many experiments, the 187 compositional boundary layers are also enriched in silica, potassium, and sulfur from dissolution of quartz, micas, and pyrite (Fig. 2a,b). These silica-rich contaminated melts are particularly significant 188 189 because in natural systems they would likely crystallise zircon, which would enable U-Pb 190 geochronology of basaltic magmas injected into shale (Gaynor *et al.*, 2022).

191 In addition to demonstrating local contamination phenomena, the experiments provide 192 compelling evidence for transfer of carbon volatiles from shale to melt. Evidence of a vapor phase is 193 preserved in the experiments as bubbles that coalesce and migrate within minutes. Our limited 194 number of experiments do not allow us to quantify how vesicle volume changes as a function of 195 reaction time, but the shale appears to undergo a significant volume change, with increasingly 196 smaller fragments visible over time (Fig. 1). The volatile mix generated is comprised dominantly of C, 197 CO, and CO₂ as revealed by spectroscopic analysis of small (ca. 20 μm diameter) subsurface bubbles 198 in a zero-time experiment (Fig. 3). Methane and water may also be present, but further analyses 199 would be required to verify this (cf. Capriolo *et al.*, 2021). Sulfur likewise cannot be ruled out as a 200 vapour phase although it appears to be largely sequestered into sulfides that dot the melt-shale 201 interface.

202 Thermal stresses and concentration of volatiles produced along cleavage planes in shale 203 would trigger a localized volume change, which could lead to fragmentation of shale and open up 204 new surface area for reaction, triggering a short-lived "run-away" effect with respect to volatile 205 expulsion. Since the host melt is virtually incapable of storing dissolved CO₂ at 150 MPa (the 206 solubility of CO_2 in basaltic magma at ca. 0.07 wt.%; Newman and Lowenstern, 2002), the system 207 would quickly become fluid oversaturated and most of the carbon in the system would be expelled 208 as a free vapor or fluid phase capable of hydrofracturing roof and wall rocks. Notably, 209 thermodynamic models predict that assimilation of as little as 0.6 wt.% organic matter will cause a

- 210 doubling of the total volatile load of a magma and produce CO-dominated gases (lacono-Marziano et
- 211 *al.*, 2012). We have now verified these predictions via direct analysis of bubbles (**Fig. 3**) and provide
- 212 unique empirical evidence for CO gas formation during magma-shale interaction.

213 Ground truthing the experimental results

214 There are several natural examples that mirror our experimental results, which we briefly summarise 215 here: (i) Shale xenoliths reported from Nuussuag (Greenland) and the Duluth Complex (USA) show 216 evidence for extensive reaction and element exchange with their host melt (Pedersen and Larsen, 217 2006; Samalens et al., 2017); (ii) Studies of mafic sills injected into shale in Skye (Scotland) show that 218 shale close to sill contacts displays abundant vesicles, partial melt textures with glassy streaks 219 enriched in sulfur, silica, and potassium, Fe-S mineralizations, and evidence for sulfur loss (Lindgren 220 and Parnell, 2006; Yallup et al., 2013); and (iii) Sill-shale contacts in the Canadian HALIP possess 221 sulfides with textures similar to those reported here (see **Supporting Information**). These natural 222 cases underscore both the utility of our experimental approach in understanding contact 223 phenomena and the dynamic, rapidly evolving nature of magma-sediment interaction.

224 WIDER IMPLICATIONS

225 Formation of sulfide xenomelts

226 Thermodynamic models have shown that as little as 0.1 wt.% magmatic assimilation of organic 227 matter (CH) will cause oxygen fugacity (fO_2) to decrease by more than two log-units (lacono-228 Marziano et al., 2012). Magmatic assimilation of shale therefore facilitates a reducing melt 229 environment while also transferring sulfur from shale to the host melt. The calculated sulfur 230 concentration at sulfide saturation (SCSS) is ca. 0.1 wt.% for the magmatic starting material at the P-231 T conditions of the experiments (Fortin et al., 2015). These low solubilities indicate that the mafic 232 host melt is ineffective as a carrier of sulfur at equilibrium conditions. Since only a minor amount of 233 sulfur is soluble in the host melt, the excess sulfur would form a dense, immiscible sulfide 234 "xenomelt" (Lesher, 2017, 2019). Some of the modified glasses in our experiments contain sulfur 235 exceeding the SCSS (Fig. 2a), which is not surprising since the experiments clearly did not attain 236 equilibrium. We therefore propose that crustally contaminated melts generated during incipient 237 magma-shale interaction can temporarily carry excess sulfur, which may manifest as a fine 238 suspension of sulfide droplets (i.e. sulfide xenomelts; Fig. 2c and Fig. 4). For the most part, the 239 experiments are visualized here in two dimensions using SEM imaging of the surface. However, 240 inspection of an experiment with extended focus under reflected light revealed a ca. 100 µm wide 241 zone around shale replete with minute (<5 μ m) sulfides beneath the sample surface (see **Supporting** 242 Information). This finding supports the notion that the compositional boundary layers in our 243 experiments are a mixture of crustally contaminated silicate melt and fine disseminations of sulfide 244 xenomelts. Our experiments therefore shed light on the origin of sulfide xenomelts in LIP systems 245 that intersect sulfur-bearing sedimentary rocks. Moreover, the experiments support the idea that 246 sulfur-charged contaminated melts can form rapidly, especially in fault-guided conduit systems 247 where magmatic and tectonic brecciation would be common (Hayes et al., 2015) and where they could be forced upward as a slug of sulfur-enriched magma. Since sulfide xenomelts can scavenge 248 249 metals from mafic magmas, they may eventually lead to the formation of economic ore deposits.

250 Climate impact

251 Our experiments highlight that magma-shale interaction in the shallow plumbing systems beneath 252 LIPs is an effective means to rapidly mobilize carbon volatiles (Fig. 4). Several major LIPs are 253 characterized by pulsed emplacement of sills into sedimentary basins (e.g. Bédard et al., 2021a, 254 2021b; Callegaro et al., 2021), which would lead to repeated episodes of magma-sediment 255 interaction and pulsed volatile release throughout the basin. These volatiles could then enter the 256 sub-aerial or sub-aqueous environment via faults and/or hydrothermal vent complexes or breccia pipes, the latter of which are known from the Karoo (Svensen et al., 2007) and Siberian LIP (Svensen 257 258 et al., 2018) and potentially the Barents Sea region of the HALIP too (Polteau et al., 2016).

259 Climate warming induced by excess carbon outgassing at LIPs would cause thermal stress in 260 ecosystems and potentially trigger ocean anoxic events (OAEs), a proximal killer in some mass 261 extinction scenarios (Bond and Grasby, 2017). The HALIP was a protracted event spanning more than 262 40 myr with a major pulse of continental basaltic magmatism between 130-120 Ma and another at ca. 95 Ma (Bédard et al., 2021a, 2021b). These pulses were coincident with OAE1a at 120 Ma and 263 264 OAE2 at 95 Ma. OAE1a was accompanied by sea-surface warming of as much as 8 °C (Ando et al., 265 2008), while OAE2 is considered to be one of the most intense OAEs (Naber et al., 2020). 266 Remarkably, the magnitude of sea-surface warming associated with OAE1a parallels that associated 267 with the Paleocene-Eocene thermal maximum and the Jurassic Toarcian OAE, consistent with the 268 idea that all of these environmental crises share a causal mechanism related to massive release of 269 carbon to the atmosphere-ocean system (Ando *et al.*, 2008).

The driving force for OAE1a and OAE2 is suspected to be volcanic CO₂ outgassing (Méhay *et al.*, 2009; Midtkandal *et al.*, 2016; Naber *et al.*, 2020) but the source(s) remain uncertain. Possibilities include dissociation of methane clathrates or volcanism from the Ontong-Java LIP, the Caribbean LIP, or the HALIP (Méhay *et al.*, 2009; Midtkandal *et al.*, 2016; Naber *et al.*, 2020). The Barents Sea sill

274 complex, part of the wider HALIP, has been speculated as a source for potentially 20,000 Gt of 275 thermogenic carbon (equivalent to 175 trillion oil barrels) via thermal metamorphism of up to 276 400,000 km³ of organic-rich sediments due to repeated sill injection (Polteau *et al.*, 2016). This 277 scenario assumes that a maximum of 2 wt.% of the carbon in thermal aureoles around sills was 278 expelled (Polteau et al., 2016), which would equate to ca. 4 wt.% original TOC assuming that no 279 more than 50% original kerogen is converted to hydrocarbon. Employing a more conservative figure 280 of 1 wt.% carbon discharged (= 2% original TOC), we arrive at 10,000 Gt thermogenic carbon 281 released from the Barents Sea volcanic basin. Important to note is that these calculations do not 282 consider inorganic carbon that could be effectively mobilized by heating of carbonates. Unfortunately, estimates of carbon release by HALIP intrusions are fraught with uncertainties 283 284 regarding quantification of parameters such as thermal aureole thicknesses and stratigraphic 285 variability in TOC. However, there are striking parallels between the estimated thermogenic carbon 286 release for the Barents Sea (Polteau et al., 2016) and for the Karoo LIP (Heimdal et al., 2021).

In this study we employed on-land HALIP sills as a case study and provided experimental 287 288 constraints that help clarify the mechanisms and rates of magma-shale interaction (summarised in Fig. 289 **4**). If we consider devolatilization of a cube of our starting material shale with 500 m side length, ca. 290 16 Mt of C could be produced. Given the short timescale of shale devolatilization observed here, and that magma volume of LIPs are on the order of 10⁵ to 10⁷ km³ (Black *et al.*, 2021), it is wholly 291 292 conceivable that several thousand Gt of C could be generated through repeated episodes of magma-293 shale interaction proximal to sills in a shallow magma plumbing system. This would moreover be in 294 addition to decarbonation in the distal, lower-temperature parts of metamorphic aureoles (processes 295 which are not directly addressed by our experiments). We therefore suggest that carbon release by 296 magma-sediment interaction at various scales during emplacement of the Canadian Arctic portion of 297 the HALIP could have been similar to estimates for thermogenic carbon release from the Barents Sea 298 (up to 20,000 Gt C; Polteau et al., 2016) or the Karoo (ca. 20,500 Gt C; Heimdal et al., 2021), making 299 the HALIP a strong contender as the causal mechanism for OAE1a and OAE2. However, detailed 300 thermal modelling of the impact of widespread HALIP sill intrusion on Sverdrup Basin sediments would 301 be required to test this hypothesis. A surprising corollary of our work is that entrapment of sulfur in 302 sulfide xenomelts might act to lessen the amount of sulfur released to the atmosphere, which in turn, 303 could boost the warming effects of carbon by producing fewer climate-cooling sulfurous aerosols. This 304 may account for the general observation (cf. Bond and Grasby, 2017) of LIPs being more strongly 305 associated with climate warming, rather than cooling.

306 ACKNOWLEDGEMENTS

307 We are grateful to the GEM-II (Geo-mapping for Energy and Minerals) program of the Geological 308 Survey of Canada and to the Polar Continental Shelf Project for logistical support in the field. We also 309 thank H. Behrens for synthesizing the hydrous starting material in Hannover (Germany), M. Nazzari 310 for EPMA support at INGV Rome (Italy), M. Choquette for EPMA analyses at Université Laval (Canada), 311 and C. Szabó for access to the Raman micro spectroscopy laboratory at Eötvös Loránd University of 312 Budapest (Hungary). We are grateful to C.M. Lesher for pre-reviewing our manuscript, to J. Heinonen, 313 D. Harlov, and M. Gavrilenko for helpful formal reviews, and to T. Waight and G. Zellmer for editorial 314 handling. This project was supported by a Swedish Research Council grant (2018-04933) to FMD.

315 DATA AVAILABILITY

316 The data underlying this article are available in the article and in its online supplementary material

317 (Supporting Information PDF file and Supporting Data File in Excel format).

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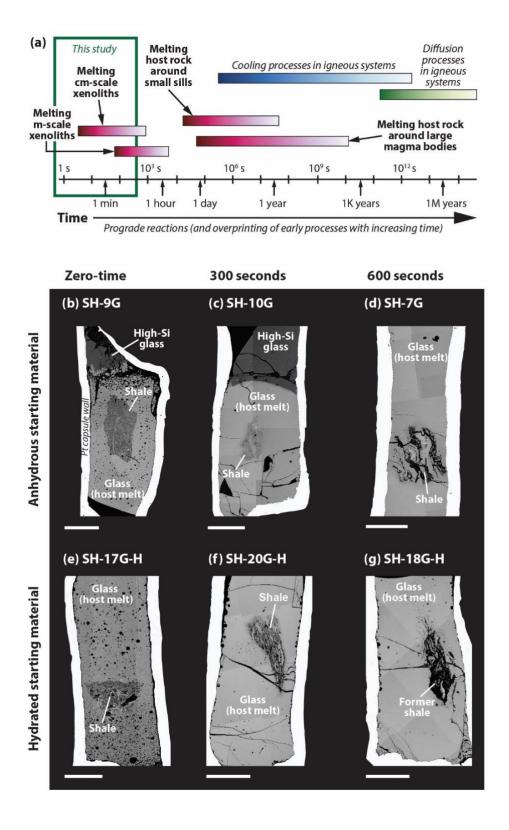




Figure 1. (a) Relative timescales for melting host rocks and xenoliths by magmatic intrusions (modified after Robertson *et al.*, 2015). This study targets the timeframe of cm- to m-scale xenolith melting in order to capture evidence of interaction processes that are only rarely preserved in nature. **(b)** to **(g)**

- 504 Backscattered electron (BSE) mosaic images of experimental run products. Intense degassing is 505 evidenced by abundant small vesicles that permeate the run products (especially in the 0 s runs). 506 Sulfides are generally found at magma-shale interfaces. Runs held for 600 s show a more advanced 507 stage of shale dissolution into the host melt. Scale bars are 1 mm.
- 508
- 509 Figure 1 caption = 106 words.
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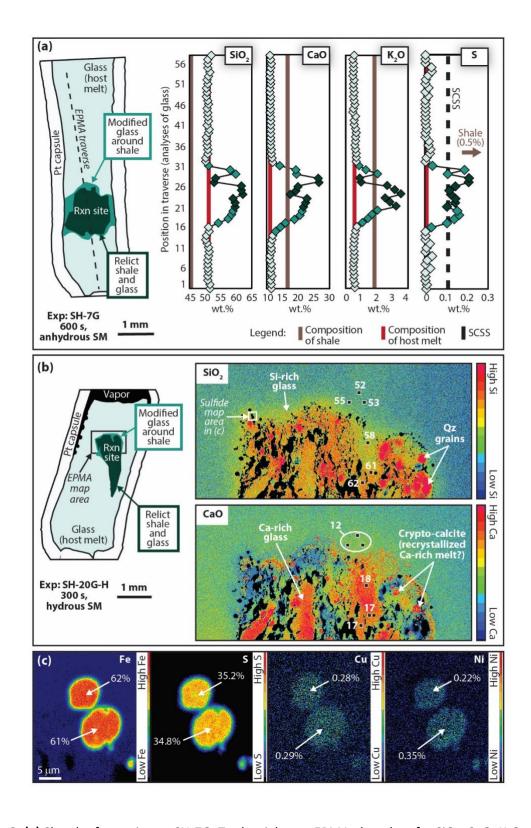
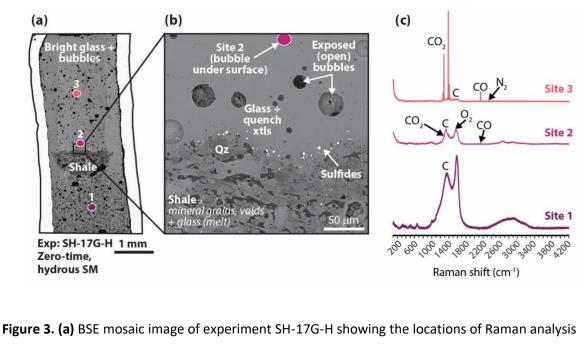


Figure 2. (a) Sketch of experiment SH-7G. To the right are EPMA glass data for SiO₂, CaO, K₂O, and S,

- respectively. (b) Sketch of experiment SH-20G-H. To the right are EPMA SiO₂ and CaO maps. Small
- 515 black squares on the maps represent EPMA glass analysis sites and numeric values are in weight %.
- 516 (c) EPMA Fe, S, Cu, and Ni maps for sulfides at the melt-shale interface (white box on SiO₂ map).
- 517 Numeric values are in weight %. Abbreviations: Qz, quartz; Rxn, reaction; SCSS, sulfur concentration

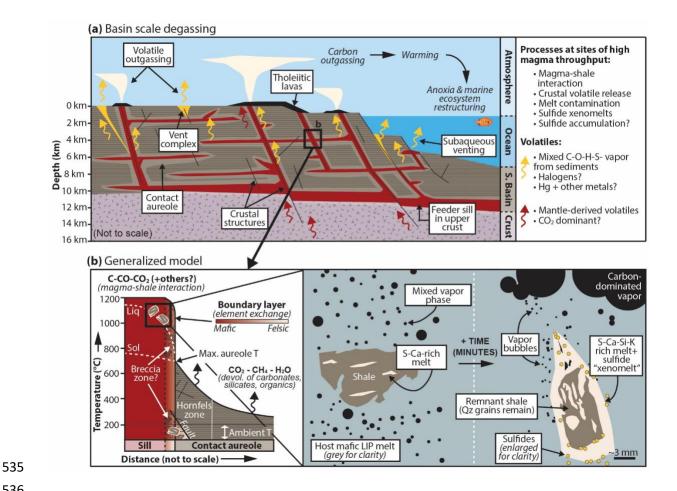
- 518 at sulfide saturation; S.M., starting material. All EPMA glass data are normalized to 100% and the
- 519 glass and sulfide data are reported in **the Supplementary Data File**.

521 Figure 2 caption = 113 words.



sites. (b) BSE image of the shale-melt interface showing the texture of partly dissolved shale, sulfides
along the shale-melt boundary, and abundant bubbles in the glass. Note that this is a surface image
and therefore the bubbles that are visible were opened during sample polishing. Raman analyses
were conducted on in-tact bubbles beneath the sample surface. (c) Raman spectra for the analysis
sites are shown to the right, with C, CO, CO₂, O₂, and N₂ bands identified (see Supplementary
Information for further details).

533 Figure 3 caption = 98 words.



537 Figure 4. (a) Illustration of a LIP plumbing system intersecting a sedimentary basin. Sill injection into 538 shale generates sediment-derived volatiles (fluids and/or vapour). If these volatiles enter the 539 atmosphere and/or ocean through a vent complex or eruption, they can trigger climate warming and 540 ocean anoxia. (b) Generalized model of sill-host rock interaction with a close-up of incipient magmashale interaction as revealed by experiments. Magma-shale interaction causes formation of carbon 541 volatiles, crustally contaminated melts, and sulfide xenomelts. Note that contact aureoles reach a 542 543 maximum temperature (max. aureole T) of ~650°C depending on multiple factors including sill temperature, ambient temperature, sill thickness, and host rock type and are a source of volatiles 544 545 not directly studied in our experiments. Sill liquidus and solidus curves are estimated based on HALIP petrological models (Bédard et al., 2021a). Abbreviations: devol, devolatilization; Liq., liquidus; Sol, 546 547 solidus; S. Basin, sedimentary basin; T, temperature.

548

Figure 4 caption = 145 words. 549